

Einstein Metrics,
Harmonic Forms, &
Conformally Kähler Geometry

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Stony Brook University

Seminario di Geometria
Terza Università di Roma
23 maggio, 2019

Dedicato

Dedicato al mio
allievo, collega, ed amico

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allievo, collega, ed amico
Massimiliano Pontecorvo,

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in occasione del convegno
MAX LX.

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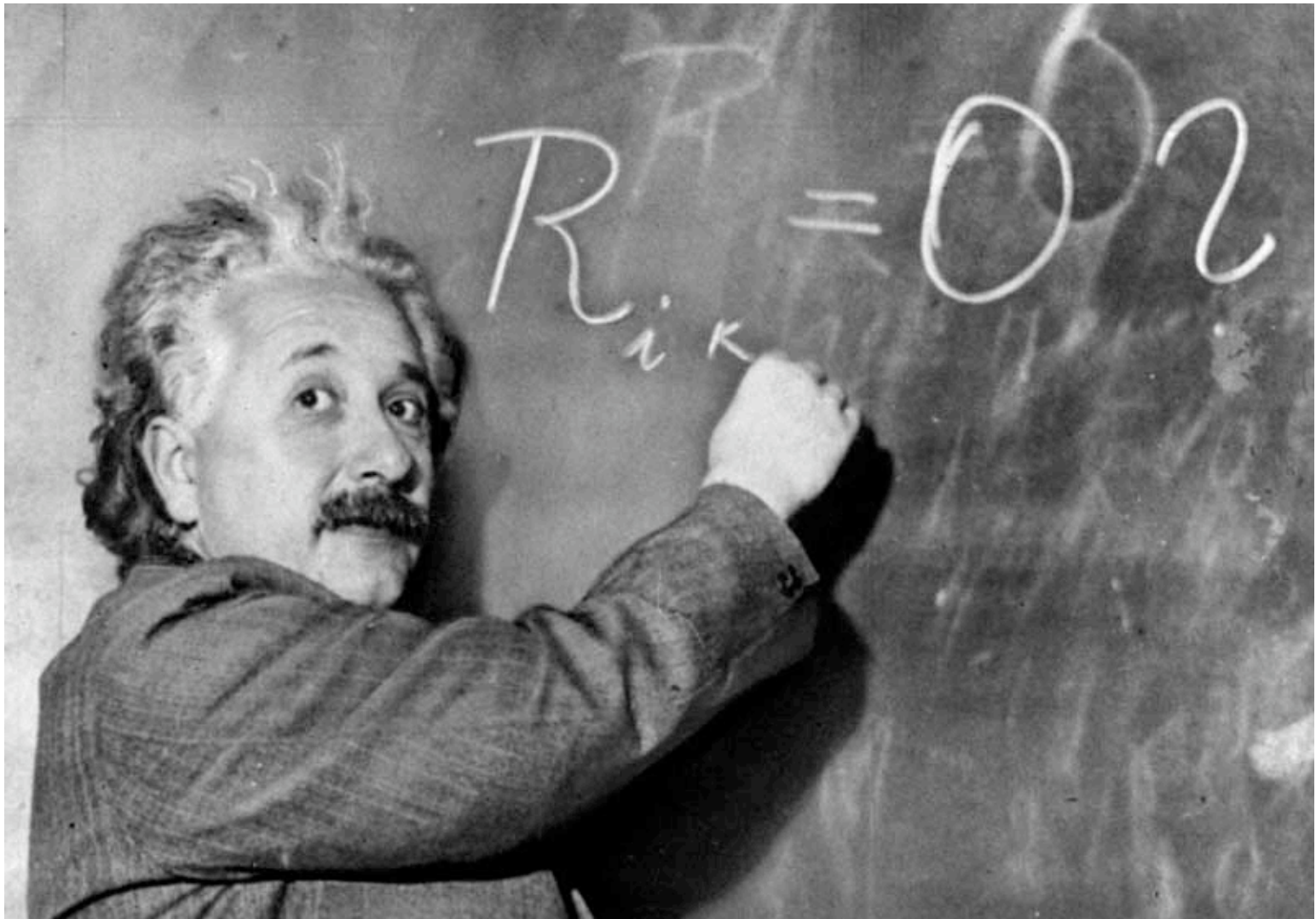
“... the greatest blunder of my life!”

— A. Einstein, to G. Gamow

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Has same sign as the *scalar curvature*

$$s = r_j^j = \mathcal{R}^{ij}_{ij}.$$

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When $n = 4$, situation is more encouraging...

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One key question:

Does enough rigidity really hold in dimension four to make this a genuine geometrization?

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Some Suggestive Questions. *If (M^4, ω) is a symplectic 4-manifold, when does M^4 admit an Einstein metric h (unrelated to ω)? What if we also require $\lambda \geq 0$?*

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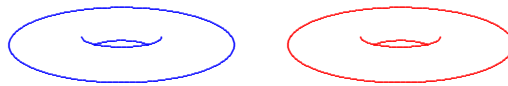
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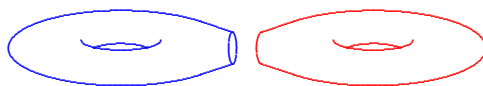
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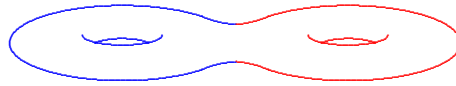
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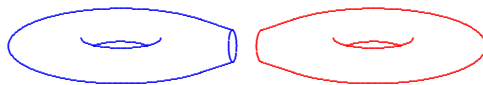
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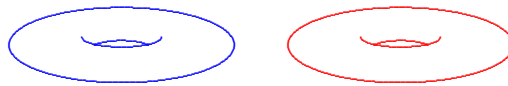
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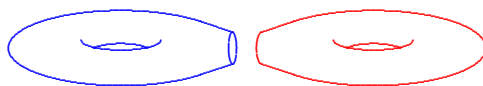
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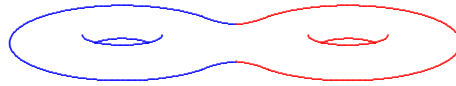
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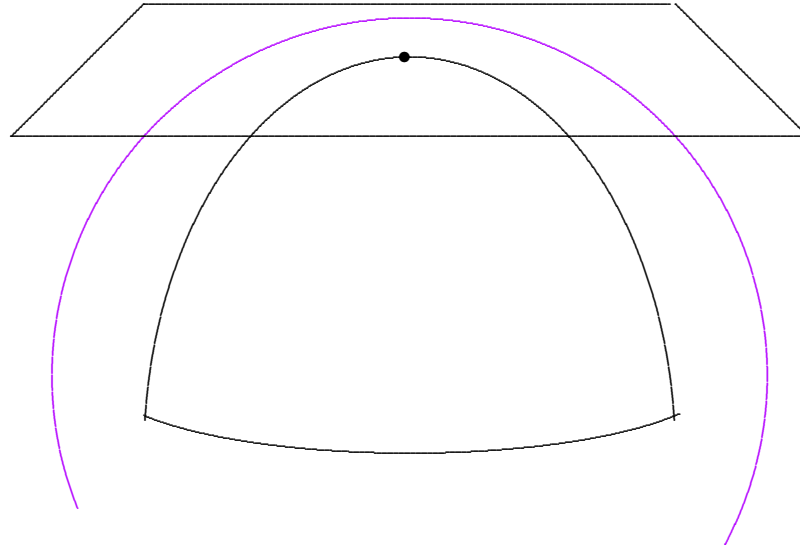
Calabi/Yau: Admits Ricci-flat Kähler metrics.

(M^n, g) :

holonomy

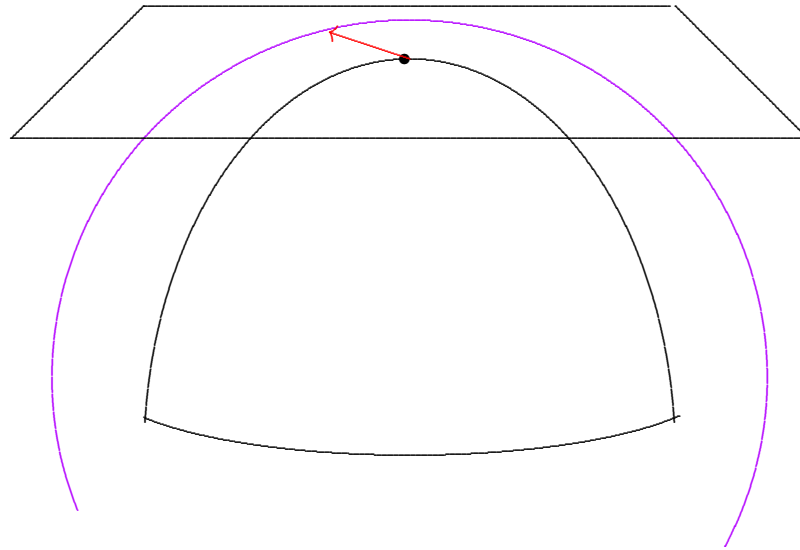
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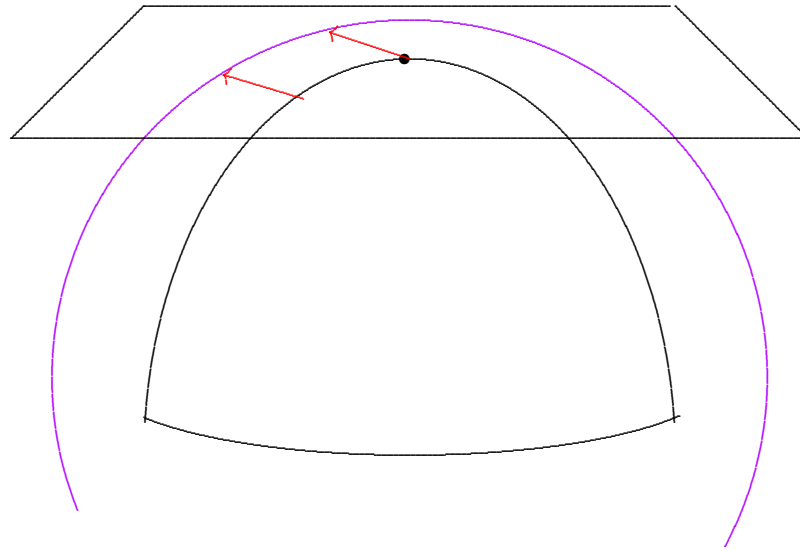
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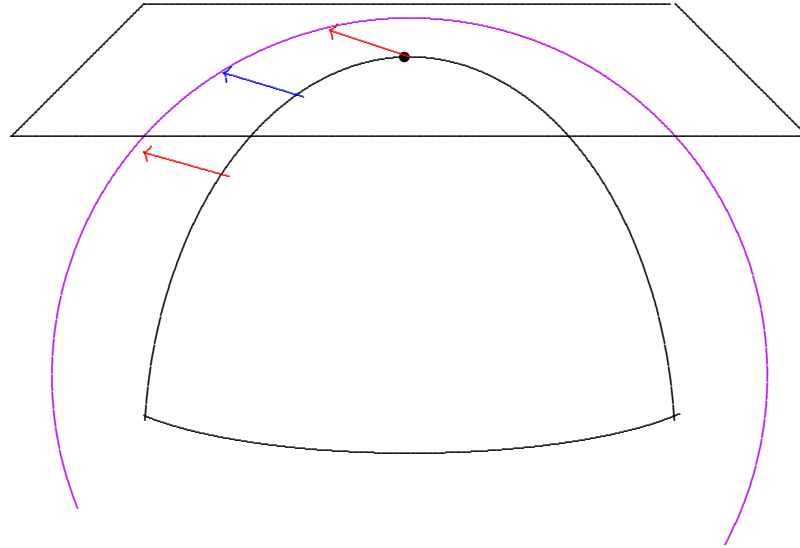
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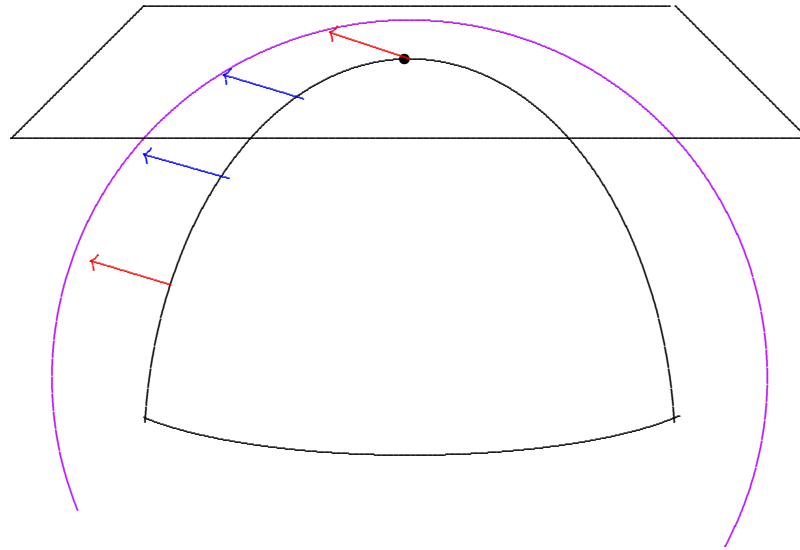
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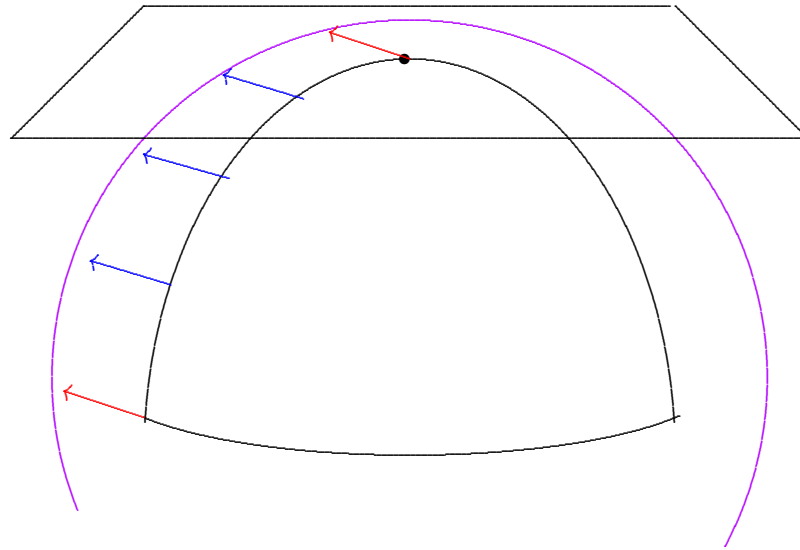
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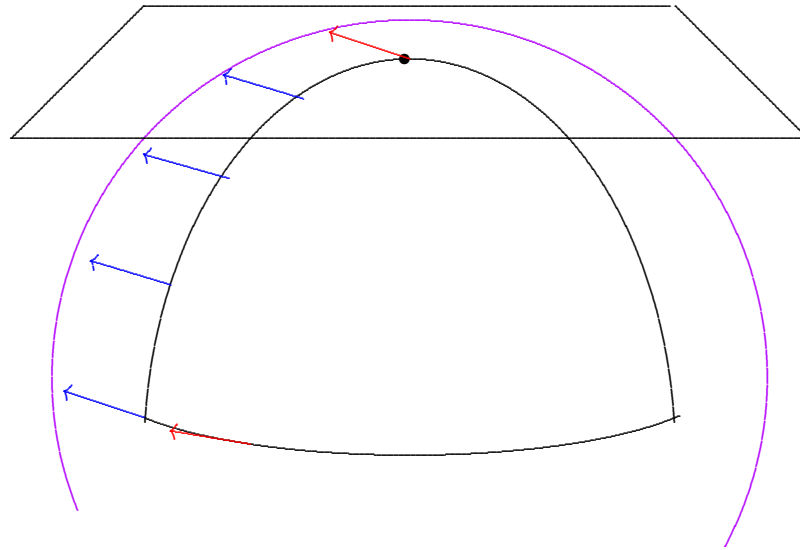
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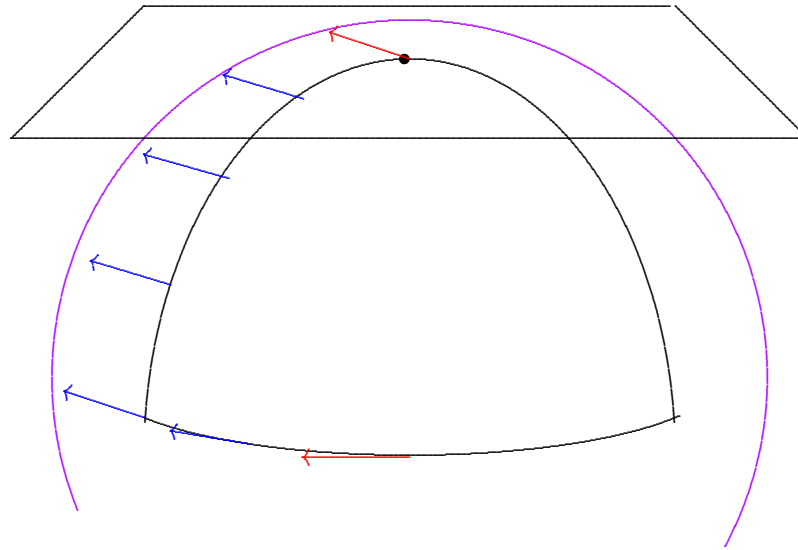
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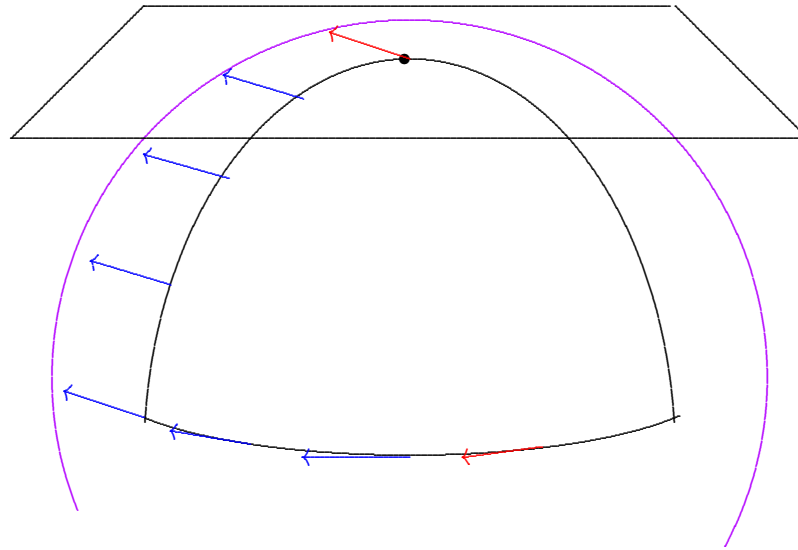
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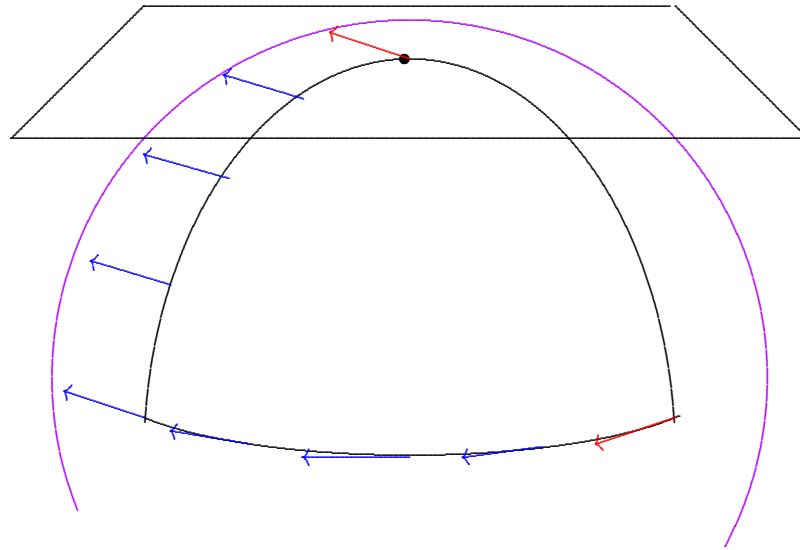
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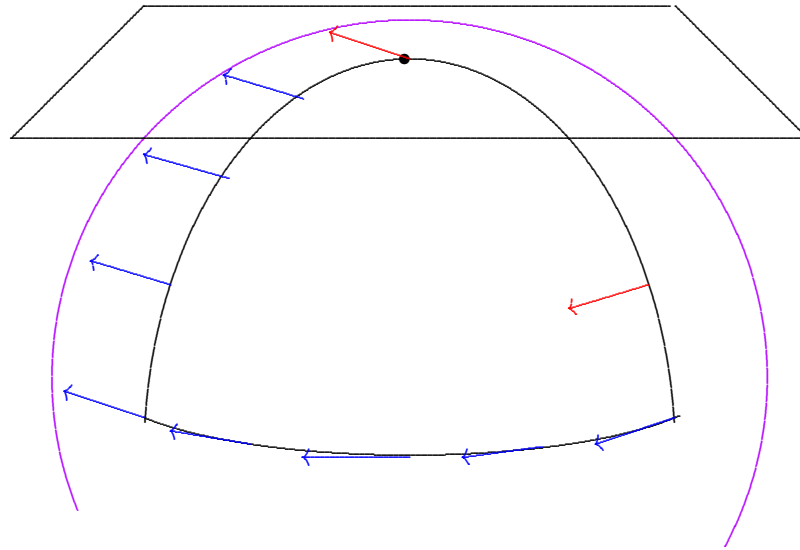
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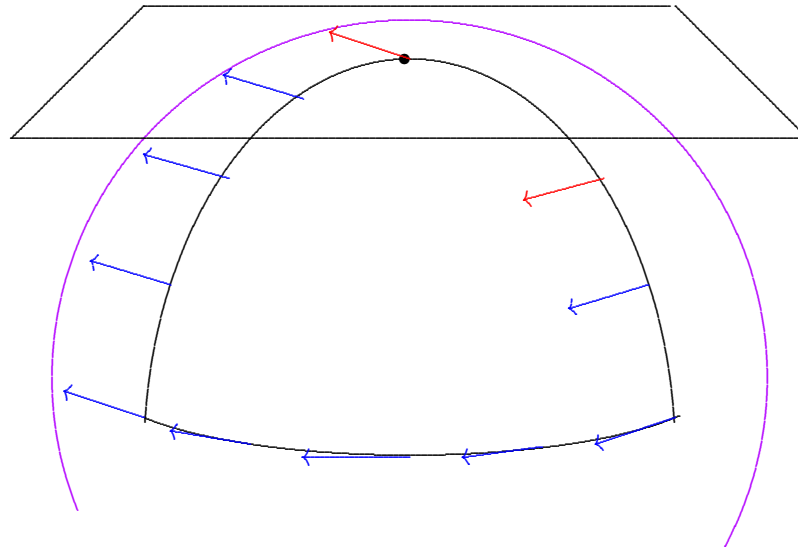
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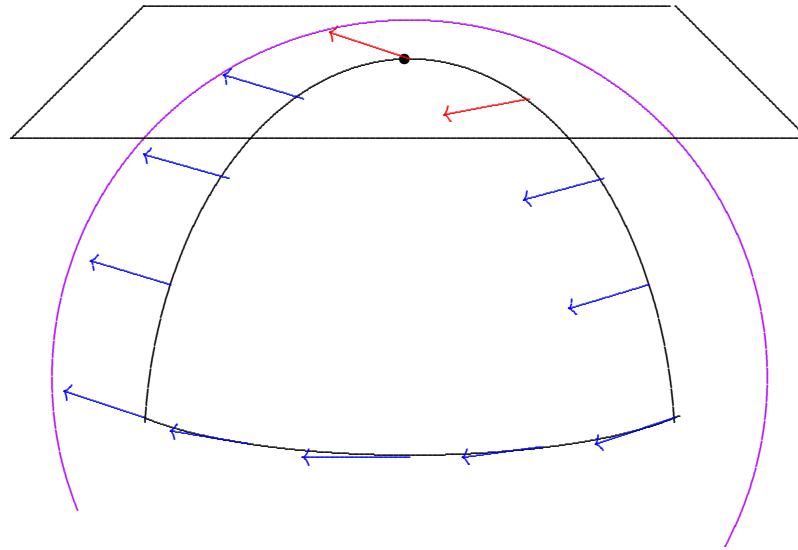
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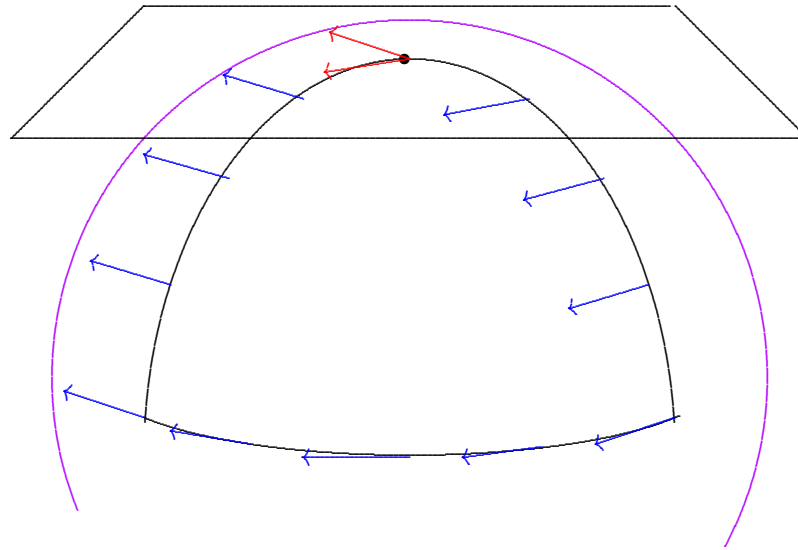
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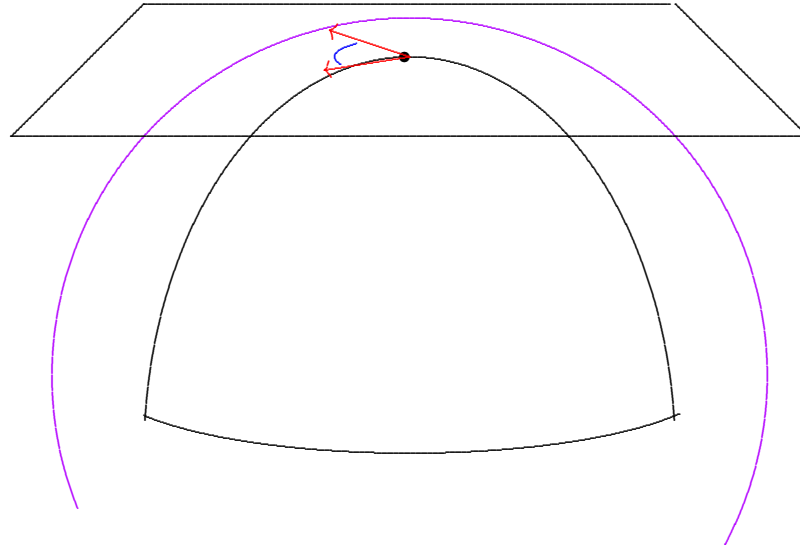
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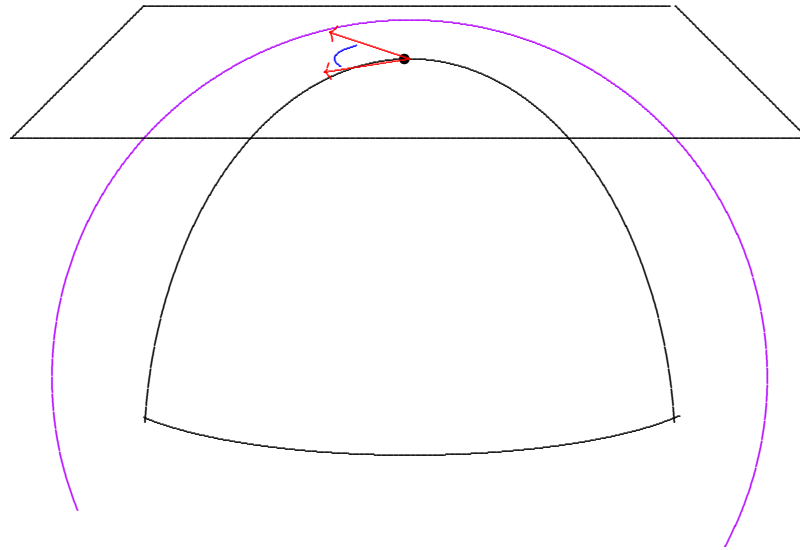
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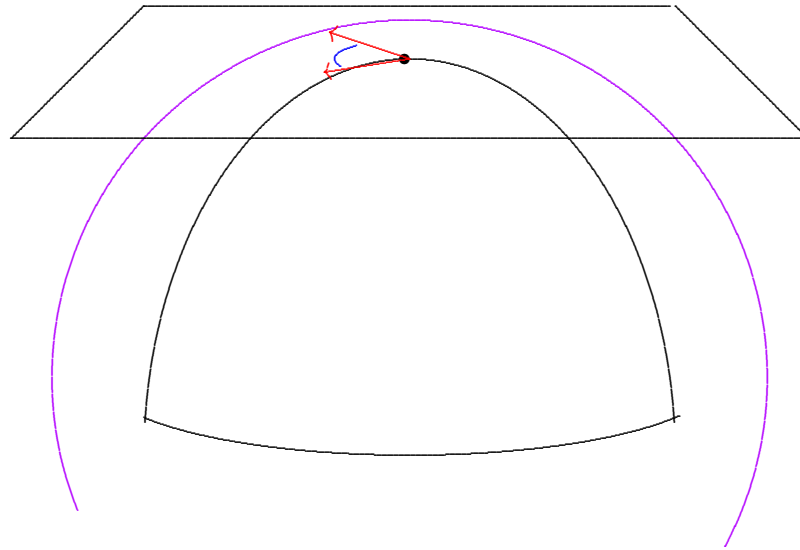
holonomy $\subset \mathbf{O}(n)$



Kähler metrics:

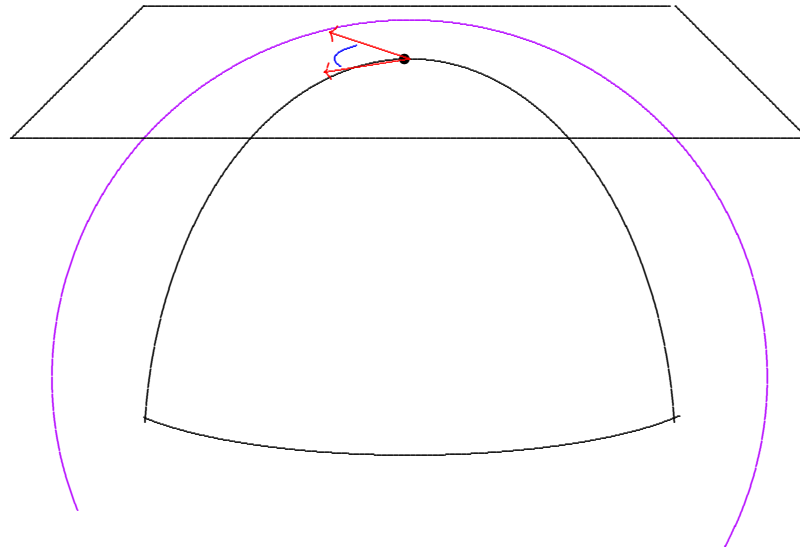
(M^{2m}, g) :

holonomy



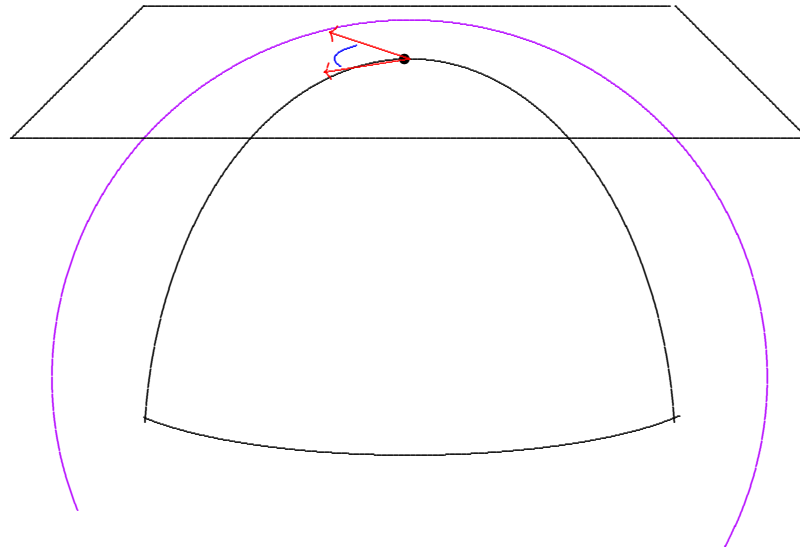
Kähler metrics:

(M^{2m}, g) Kähler \iff holonomy $\subset \mathbf{U}(m)$



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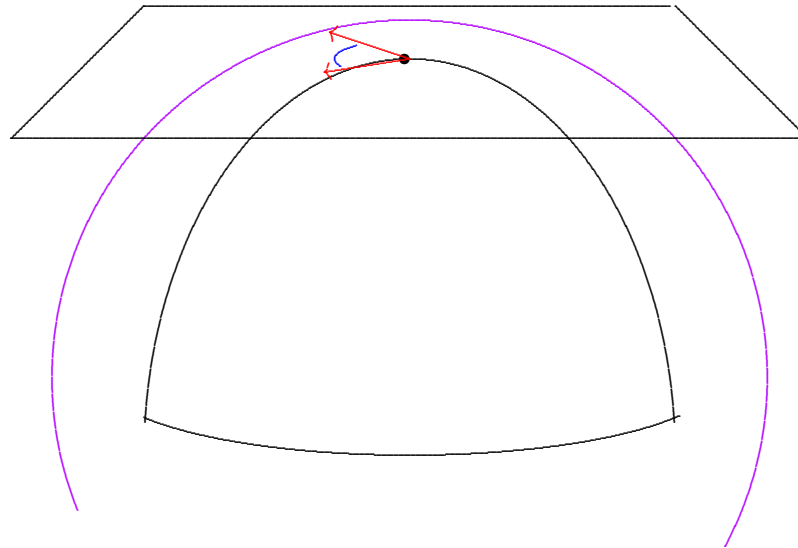
(M^{2m}, g) Kähler \iff holonomy $\subset \mathbf{U}(m)$



$$\mathbf{U}(m) := \mathbf{O}(2m) \cap \mathbf{GL}(m, \mathbb{C})$$

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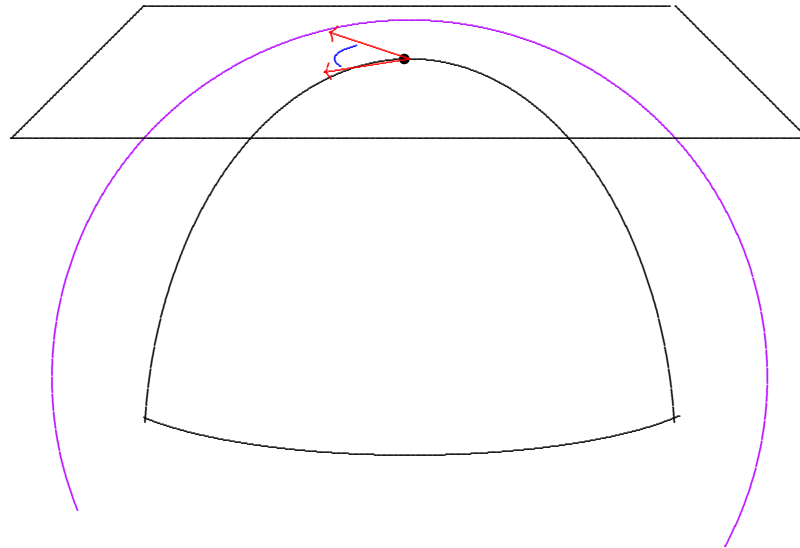
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Makes tangent space a complex vector space!

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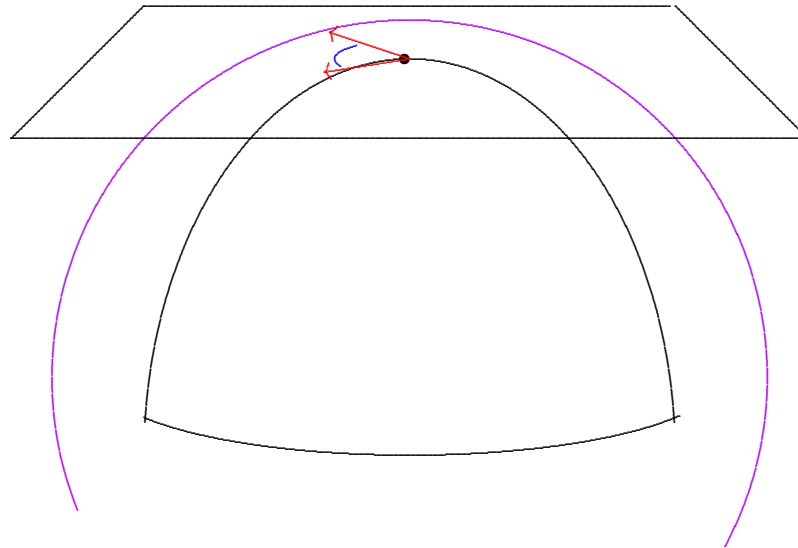
Makes tangent space a complex vector space!

$$J : TM \rightarrow TM, \quad J^2 = -\text{identity}$$

“almost-complex structure”

Kähler metrics:

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Makes tangent space a complex vector space!

Invariant under parallel transport!

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$\iff \exists$ almost complex-structure J with $\nabla J = 0$
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ω called “Kähler form.”

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$$d\omega = 0$$

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Kähler magic:

If we define the Ricci form by

$$\rho = r(J\cdot, \cdot)$$

then $i\rho$ is curvature of canonical line bundle $\Lambda^{m,0}$.

Kähler metrics:

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$$[\omega] \in H^2(M)$$

“Kähler class”

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ω non-degenerate closed 2-form:

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ω non-degenerate closed 2-form: symplectic form

Theorem (L '09). *Suppose that M is a smooth compact oriented 4-manifold which admits a symplectic structure ω . Then M also admits an Einstein metric h with $\lambda \geq 0$ if and only if*

$$M \stackrel{\text{diff}}{\approx} \left\{ \begin{array}{l} \mathbb{C}P_2 \# k \overline{\mathbb{C}P}_2, \quad 0 \leq k \leq 8, \\ S^2 \times S^2, \\ K3, \\ K3/\mathbb{Z}_2, \\ T^4, \\ T^4/\mathbb{Z}_2, T^4/\mathbb{Z}_3, T^4/\mathbb{Z}_4, T^4/\mathbb{Z}_6, \\ T^4/(\mathbb{Z}_2 \oplus \mathbb{Z}_2), T^4/(\mathbb{Z}_3 \oplus \mathbb{Z}_3), \text{ or } T^4/(\mathbb{Z}_2 \oplus \mathbb{Z}_4). \end{array} \right.$$

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K3 surface, Enriques surface,

Abelian surface, Hyper-elliptic surfaces.

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Definitive list ...

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Every Einstein metric is Ricci-flat Kähler.

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Moduli space $\mathcal{E}(M)$

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Moduli space $\mathcal{E}(M) = \{\text{Einstein } h\} / (\text{Diffeos} \times \mathbb{R}^+)$

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Moduli space $\mathcal{E}(M)$ completely understood.

But we understand some cases better than others!

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Below the line:

Every Einstein metric is Ricci-flat Kähler.

Moduli space $\mathcal{E}(M)$ connected!

Above the line:

Know an Einstein metric on each manifold.

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Below the line:

Every Einstein metric is Ricci-flat Kähler.

Moduli space $\mathcal{E}(M)$ connected!

Above the line:

Moduli space $\mathcal{E}(M) \neq \emptyset$.

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Below the line:

Every Einstein metric is Ricci-flat Kähler.

Moduli space $\mathcal{E}(M)$ connected!

Above the line:

Moduli space $\mathcal{E}(M) \neq \emptyset$. But is it connected?

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Del Pezzo surfaces:

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(M^4, J) for which c_1 is a Kähler class $[\omega]$.

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Blow-up of $\mathbb{C}P_2$ at k distinct points,
in general position,

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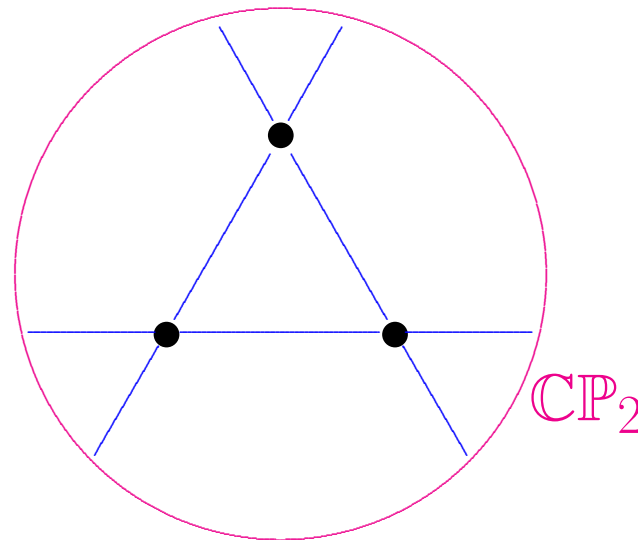
Blow-up of $\mathbb{C}P_2$ at k distinct points, $0 \leq k \leq 8$,
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Blowing up:

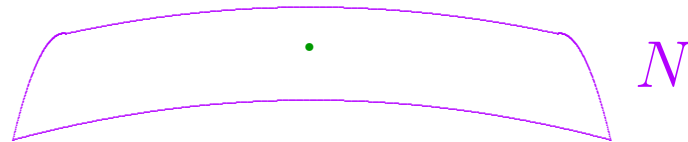
Blowing up:

If N is a complex surface,



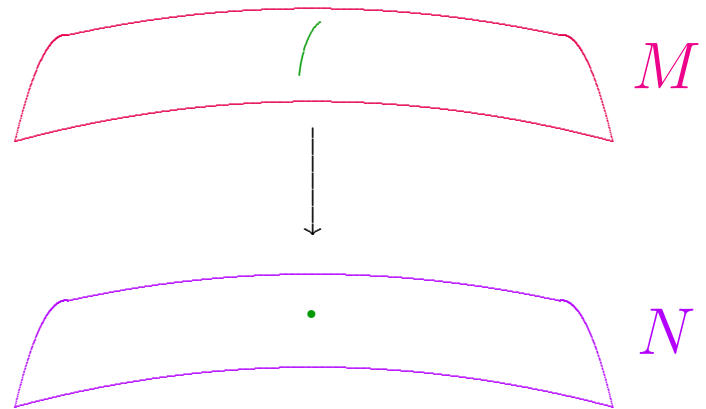
Blowing up:

If N is a complex surface, may replace $p \in N$



Blowing up:

If N is a complex surface, may replace $p \in N$
with $\mathbb{C}P_1$

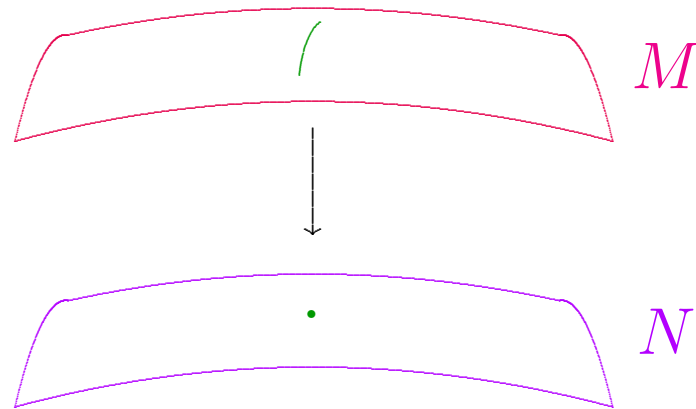


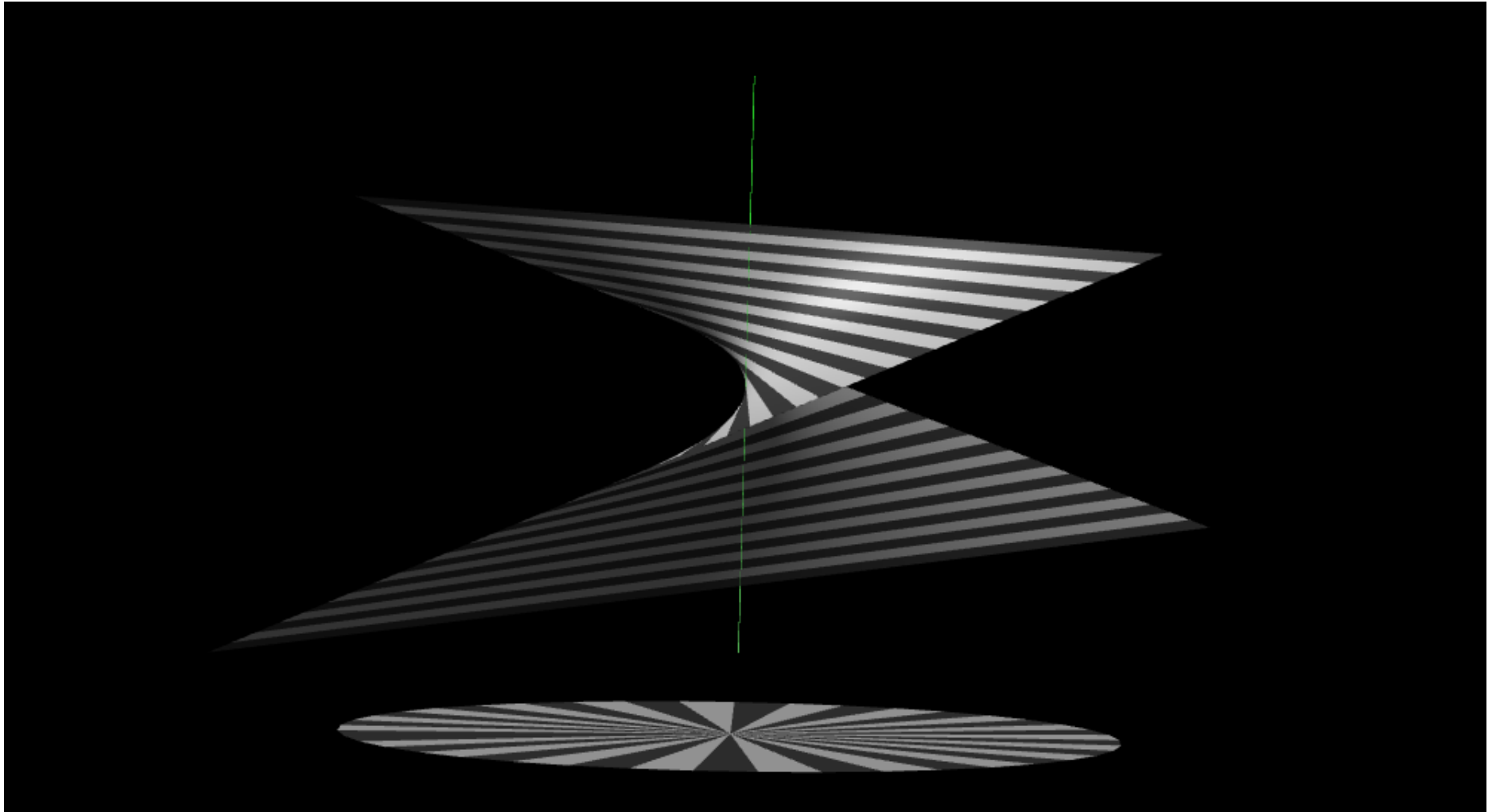
Blowing up:

If N is a complex surface, may replace $p \in N$ with $\mathbb{C}P_1$ to obtain **blow-up**

$$M \approx N \# \overline{\mathbb{C}P_2}$$

in which added $\mathbb{C}P_1$ has normal bundle $\mathcal{O}(-1)$.



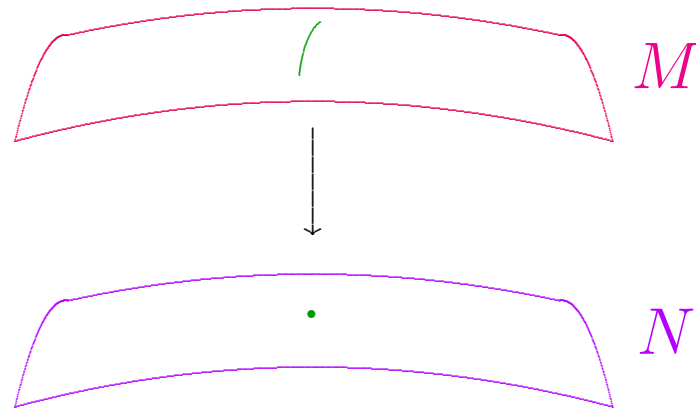


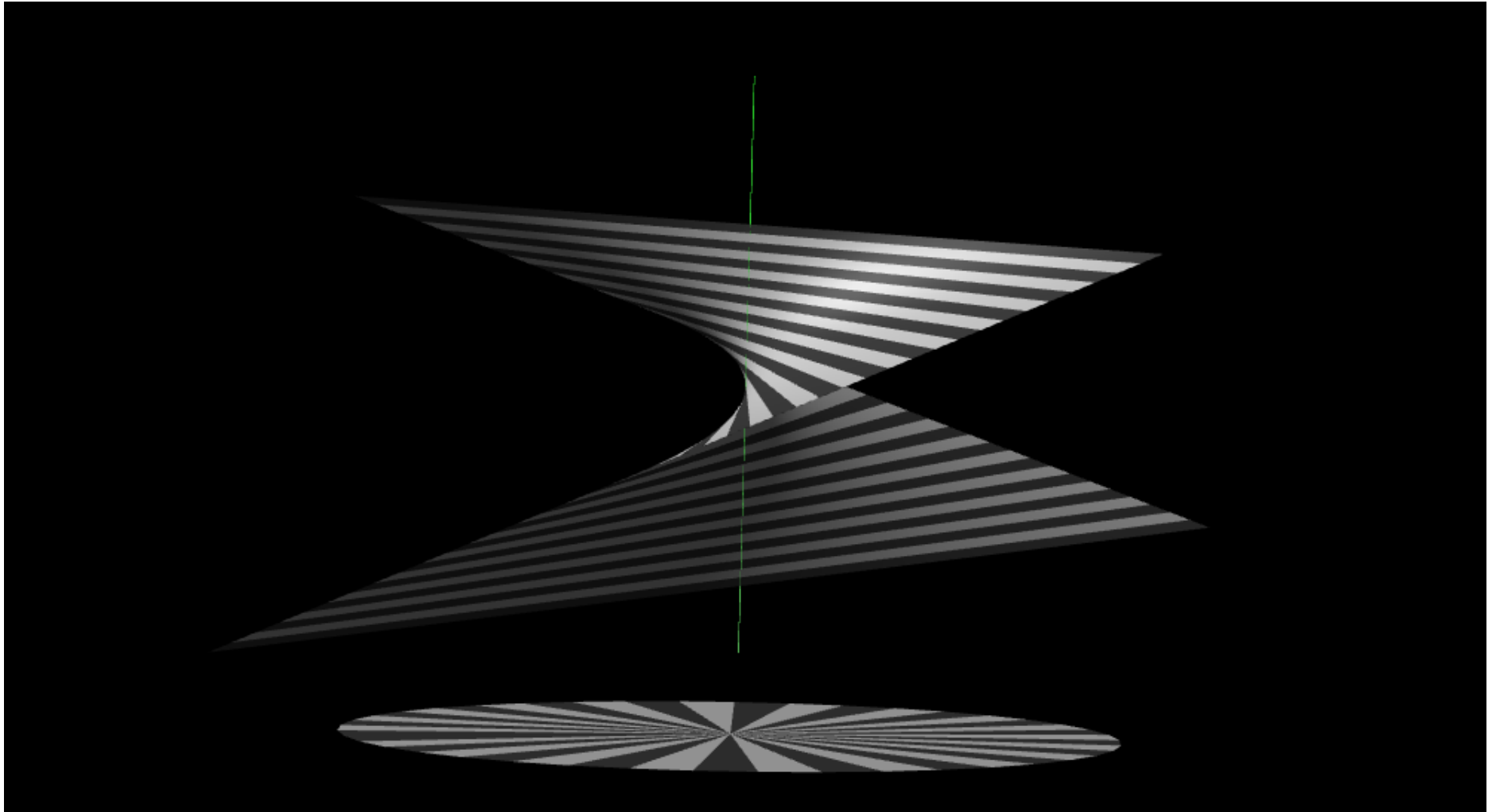
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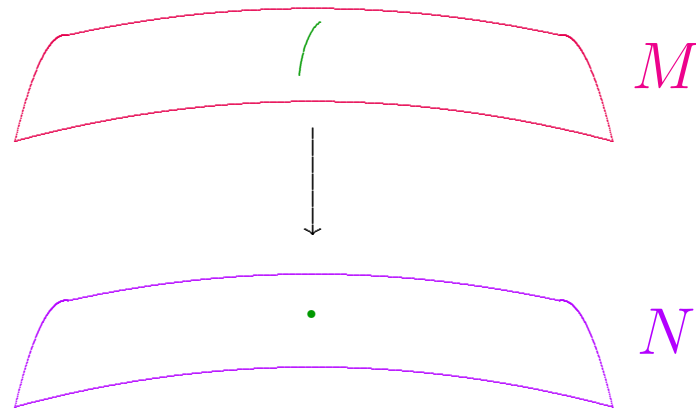


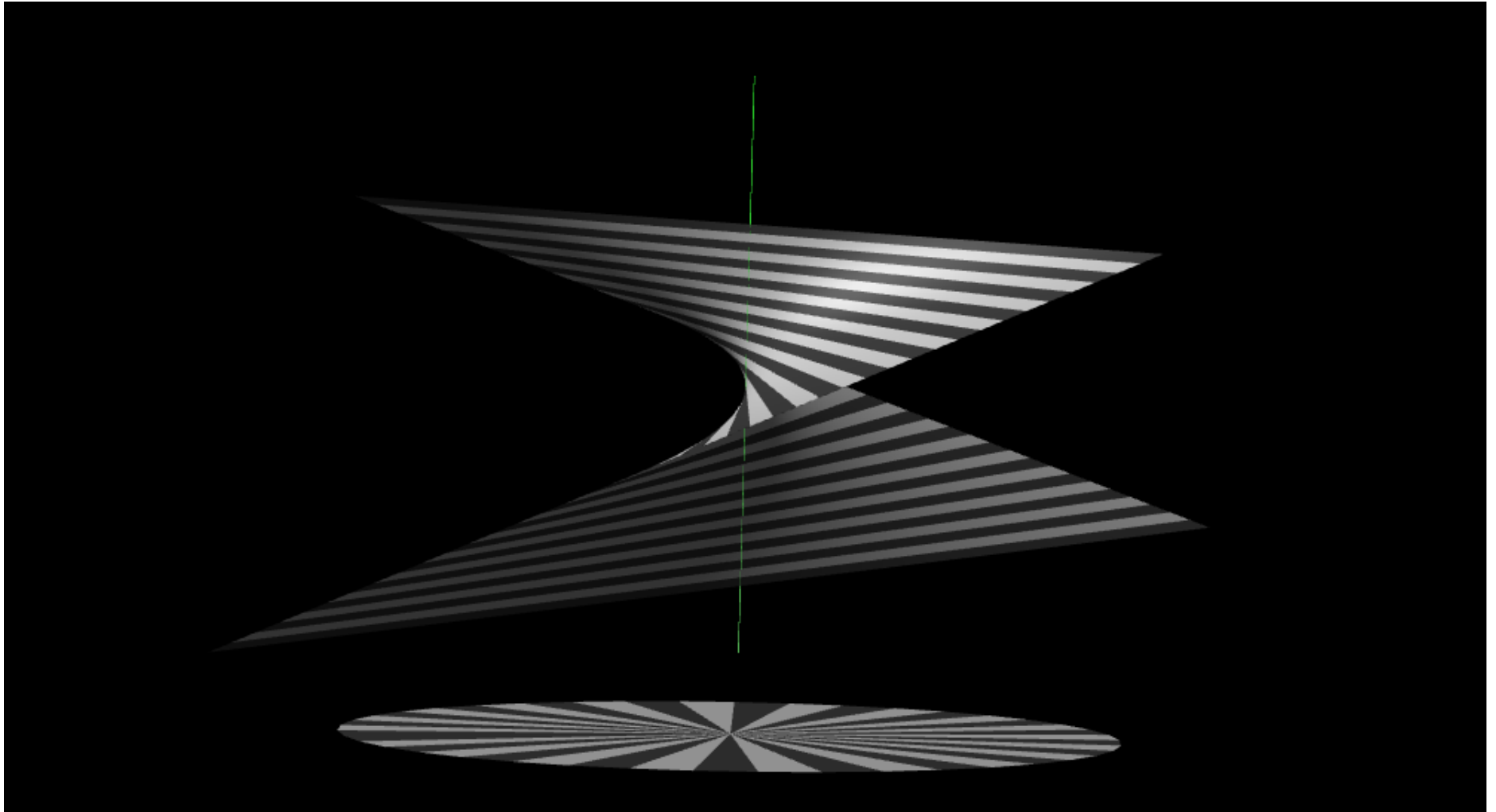
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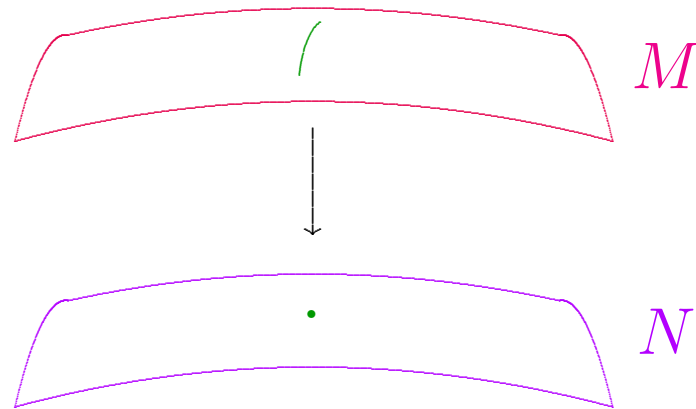


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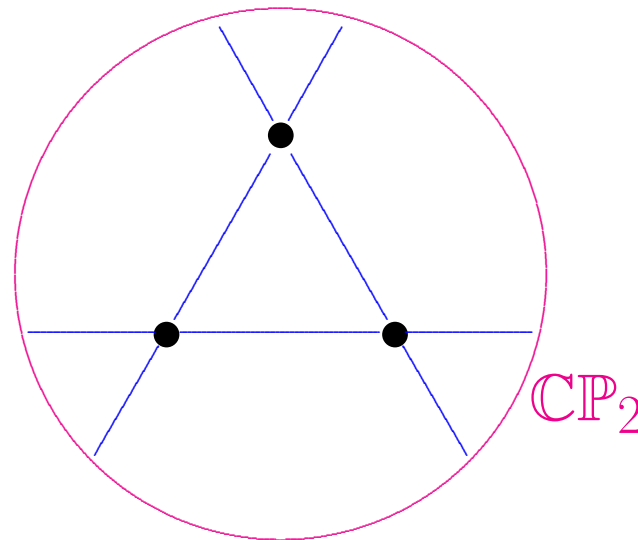


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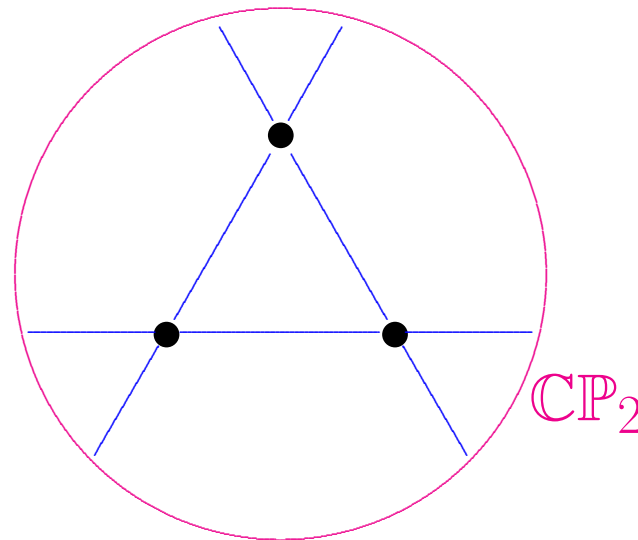
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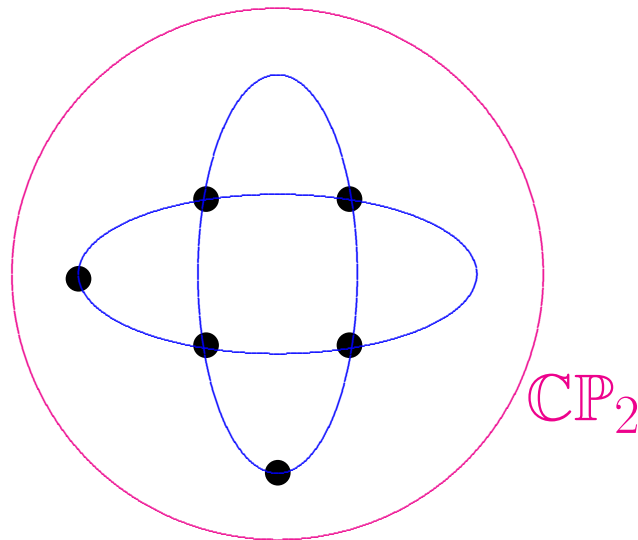


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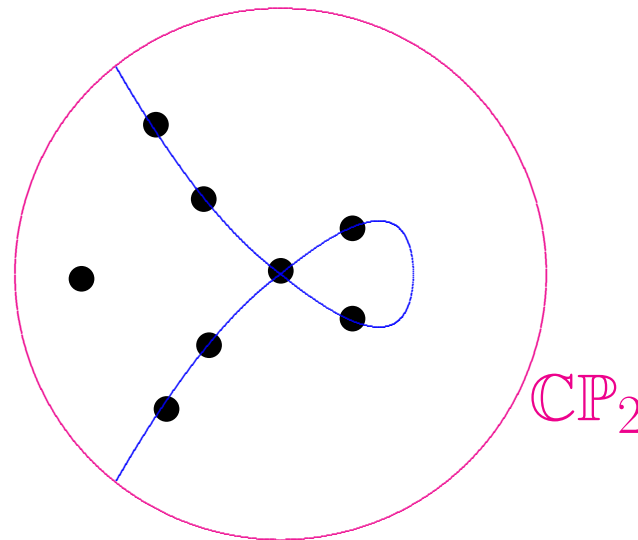


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Uniqueness: Bando-Mabuchi '87, L '12.

Above the line:

Moduli space $\mathcal{E}(M) \neq \emptyset$. But is it connected?

$$\mathbb{C}P_2 \# k \overline{\mathbb{C}P_2}, \quad 0 \leq k \leq 8,$$
$$S^2 \times S^2,$$

$K3$,

$K3/\mathbb{Z}_2$,

T^4 ,

$T^4/\mathbb{Z}_2, T^4/\mathbb{Z}_3, T^4/\mathbb{Z}_4, T^4/\mathbb{Z}_6$,

$T^4/(\mathbb{Z}_2 \oplus \mathbb{Z}_2), T^4/(\mathbb{Z}_3 \oplus \mathbb{Z}_3)$, or $T^4/(\mathbb{Z}_2 \oplus \mathbb{Z}_4)$.

Below the line:

Every Einstein metric is Ricci-flat Kähler.

Moduli space $\mathcal{E}(M)$ connected!

Basic problem:

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Understand all Einstein metrics on Del Pezzos.

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More generally, their dimensions

$$b_\pm(M) = \dim \mathcal{H}_h^\pm$$

are completely metric-independent, and are oriented homotopy invariants of M .

Key background result:

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Yes — with a reasonable extra hypothesis on ω ...

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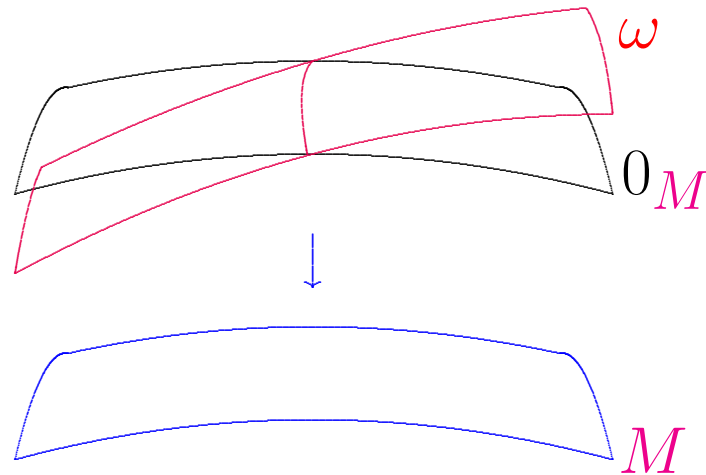
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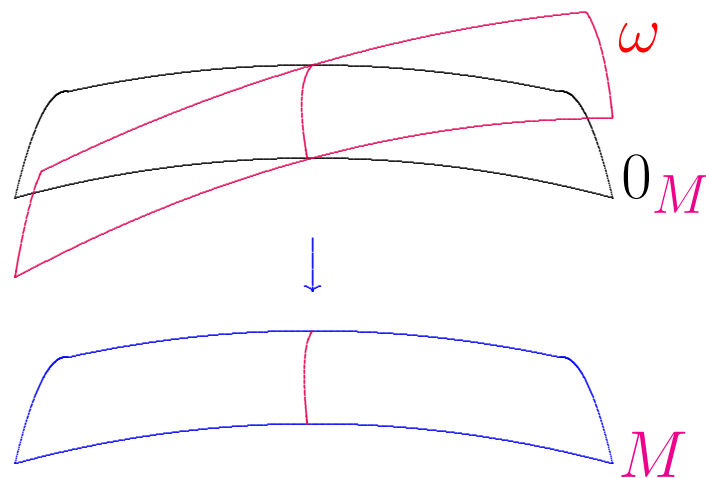
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\implies Zero set Z of ω has codimension 3:

$$Z \approx \sqcup_{j=1}^n S^1.$$

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Theorem (Taubes, et al). If $b_+(M) \neq 0$, such forms exist for an open dense set of metrics h on M .

Theorem A.

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Indeed, all **known** Einstein metrics on Del Pezzo surfaces arise this way!

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Conversely, these complex surfaces all admit $\lambda \geq 0$ Einstein metrics h of the above type.

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Moral: Taubes' genericity result does not guarantee genericity among metrics solving an equation!

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$$W^+ = \begin{bmatrix} -\frac{s}{12} & & \\ & -\frac{s}{12} & \\ & & \frac{s}{6} \end{bmatrix}$$

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$$W^+(\omega, \omega) \geq 0, \quad W^+(\omega, \omega) \neq 0$$

Before discussing **Theorems A & B**,
consider simpler case when $W^+(\omega, \omega) > 0$.

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for some g -preserving almost-complex structure J .

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as proxy for Einstein equation.

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for $fW^+ \in \text{End}(\Lambda^+)$.

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(for any almost-Kähler 4-manifold)

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where $W^+(\omega)^\perp =$ projection of $W^+(\omega, \cdot)$ to ω^\perp .

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This yields

$$0 \geq 3 \int_M W^+(\omega, \omega) \left(W^+(\omega, \omega) - \frac{s}{3} \right) f \, d\mu$$

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$\therefore h \propto s^{-2}g$ globally on M .

Tanti auguri, e buon compleanno, Max!

